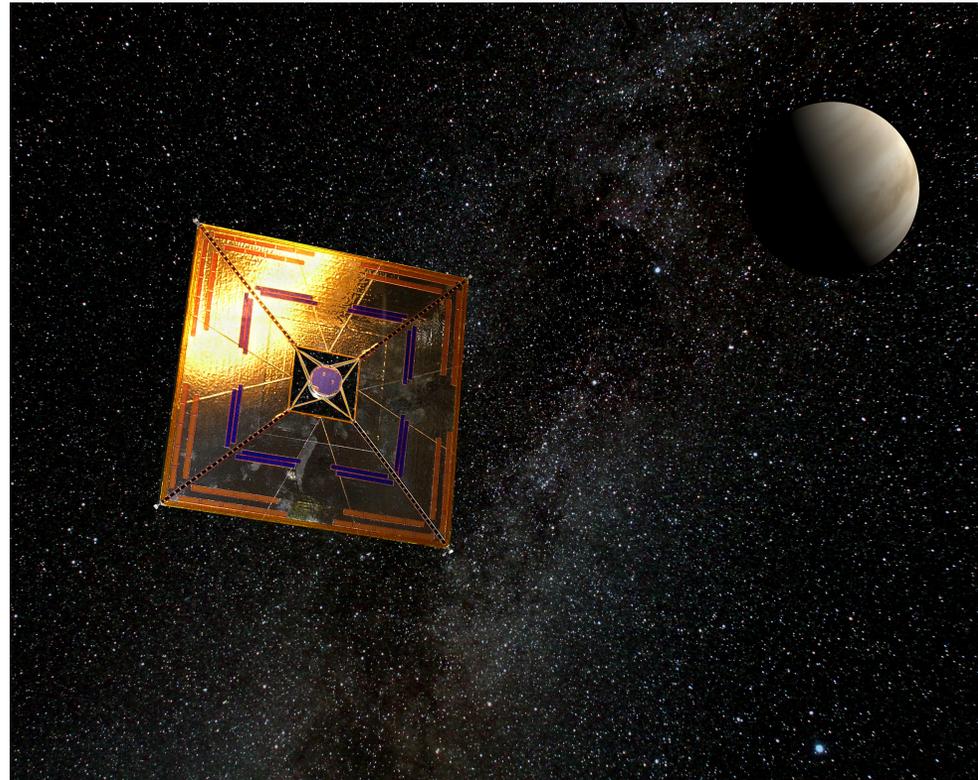


AA103

Air and Space Propulsion

Space Sailing



Properties of light

- Momentum

$$p = \frac{h}{\lambda}$$

- Energy

$$E = h\nu = h \frac{c}{\lambda} = pc \quad ; \quad \begin{array}{l} h = 6.63 \times 10^{-34} \text{ Joule} - \text{sec} \\ c = 3.00 \times 10^8 \text{ M / sec} \end{array}$$

Reference - *Space Sailing* by Jerome L. Wright, Gordon and Breach Science Publishers 1994

Much more current - https://en.wikipedia.org/wiki/Solar_sail

Properties of light, cont' d

- Energy flux

$$W = \left[\frac{\text{Joules}}{\text{photon}} \right] \cdot \left[\frac{\text{photons}}{\text{M}^2 - \text{sec}} \right] = h\nu \cdot \left[\frac{\text{photons}}{\text{M}^2 - \text{sec}} \right] = \left[\frac{\text{Joules}}{\text{M}^2 - \text{sec}} \right]$$

At the earth's radius from the sun

$$W_{\text{earth}} = 1368 \text{ Joules} / \text{M}^2 - \text{sec}$$

$$W_{\text{earth}} / c = 4.56 \times 10^{-6} \text{ N} / \text{M}^2$$

Properties of light, cont' d

- Light pressure on a perfectly reflecting surface normal to the incidence direction of light

$$P = 2W / c$$

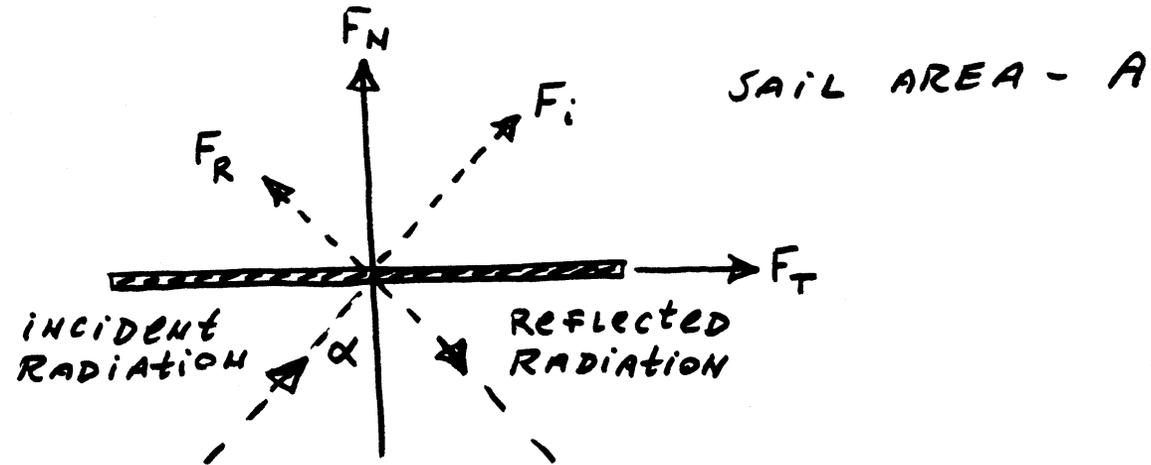
At the earth's radius

$$P_{\text{earth}} = 9.12 \times 10^{-6} \text{ N / M}^2$$

At other radii

$$P = \left(9.12 \times 10^{-6} \text{ N / M}^2 \right) \left(\frac{r_{\text{earth}}}{r} \right)^2 ; \quad \frac{r}{r_{\text{earth}}} = \text{radius in AU}$$

Light Force on a Sail



$$F_N = F_i \cos \alpha + F_R \cos \alpha \quad ; \quad F_T = F_i \sin \alpha - F_R \sin \alpha$$

- Perfect reflection

$$F_i = \frac{W}{c} A \cos \alpha \quad ; \quad F_R = \frac{W}{c} A \cos \alpha$$

$$F_N = 2 \frac{W}{c} A \cos^2 \alpha \quad ; \quad F_T = 0$$

Light Force on a Sail, cont' d

- Taking account of reflected, absorbed and radiated energy

$$\frac{F_N}{\left(2 \frac{W}{c} A\right)} = \frac{(1 + rs) \cos^2 \alpha}{2} + \frac{B_f r (1 - s) \cos \alpha}{2} + \frac{B_f e_f - B_b e_b}{e_f + e_b} \frac{(1 - r) \cos \alpha}{2}$$

$$\frac{F_T}{\left(2 \frac{W}{c} A\right)} = \frac{(1 - rs) \cos \alpha \sin \alpha}{2}$$

where

r = reflectivity of the front surface for the incident radiation

s = specular reflection coefficient

e_f, e_b = front and back surface IR emission coefficients for wavelength of emitted radiation based on sail temperature.

B_f, B_b = Non-Lambertian coefficients for front and back surfaces.

Sail acceleration

The size of a sail is determined by the mass of the payload and the characteristic acceleration required for a particular mission.

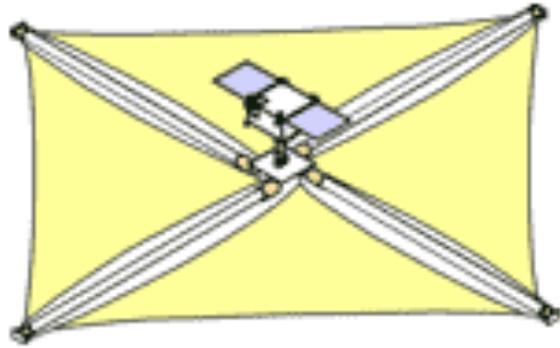
$$a_c = 2\eta \frac{W}{c} \left(\frac{A}{m_{\text{total}}} \right)$$

where m_{total} is the total mass of the ship and η is the sail efficiency (typically about 0.9). The key factor limiting the acceleration available is the mass loading of the sail.

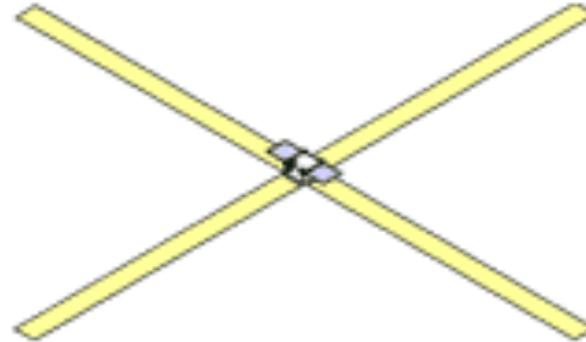
$$\sigma = \frac{m_{\text{total}}}{A}$$

The lowest available mass loading using currently available materials is about 5 gm/M²

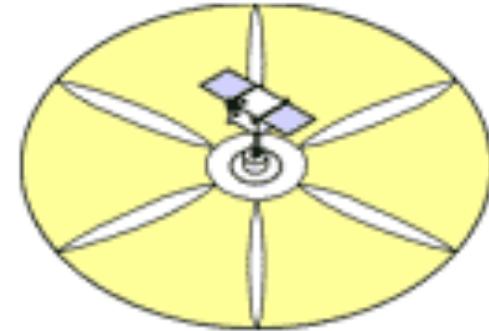
Sail design concepts



Square Sail (not to scale)



Heliogyro (not to scale)

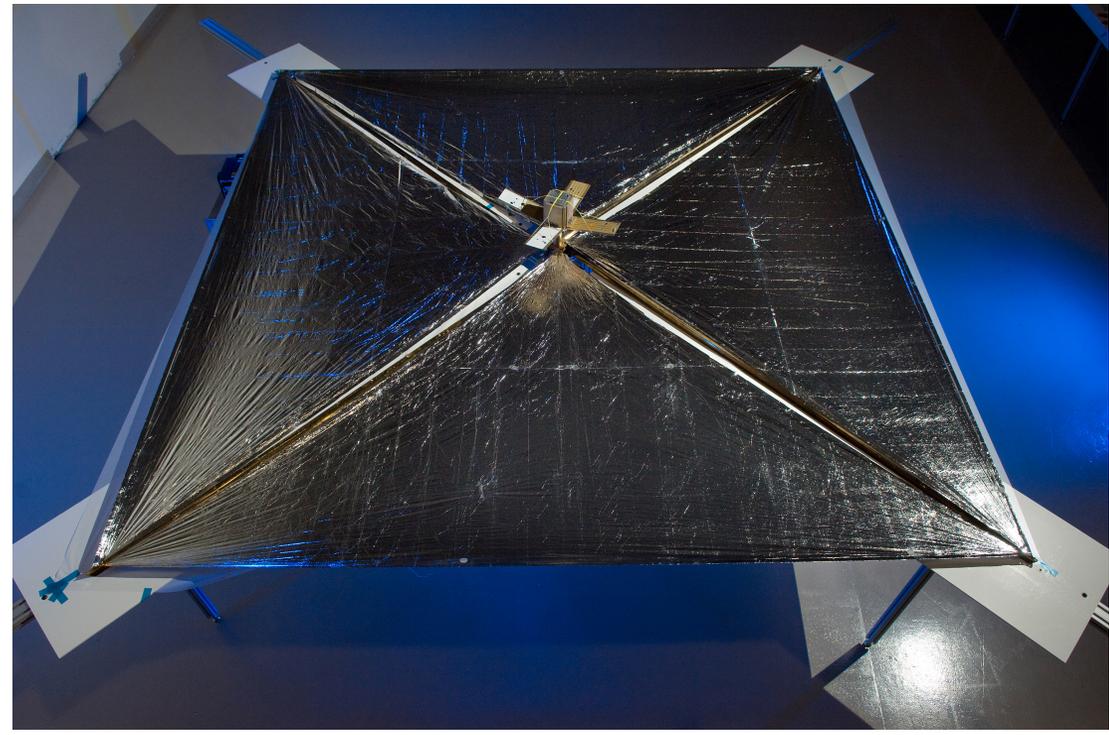
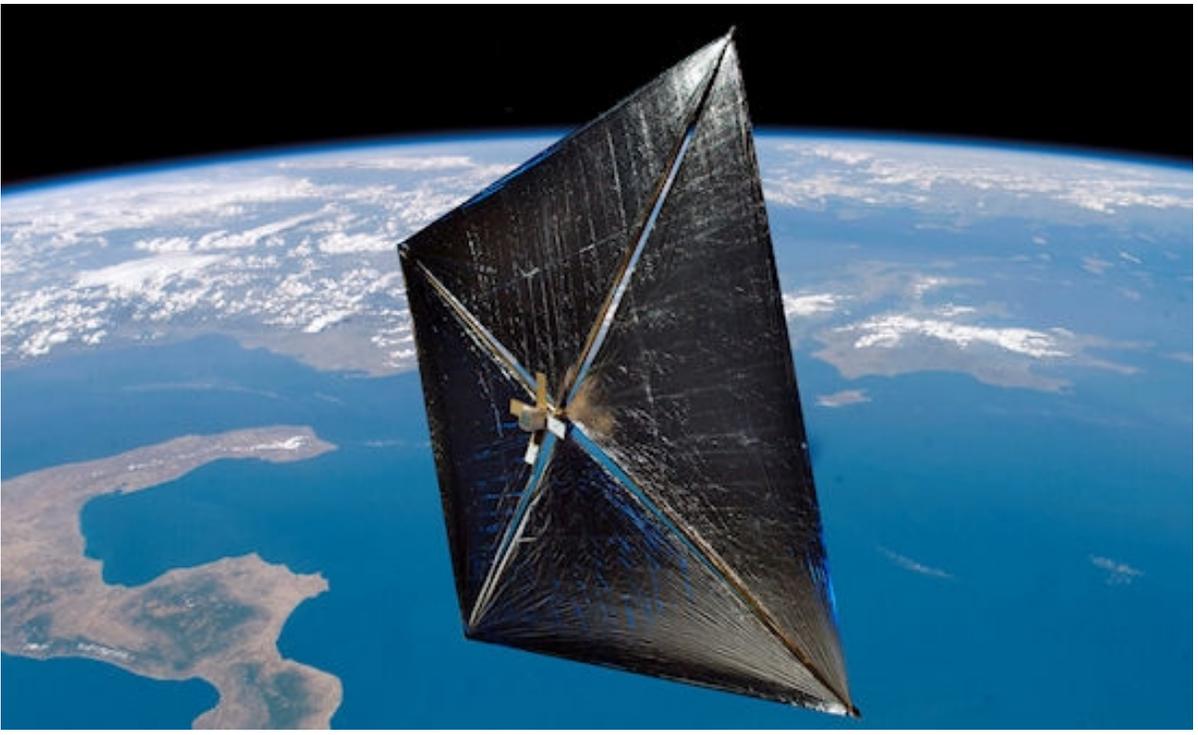


**Spinning Disk Sail
(not to scale)**

Recent successful missions

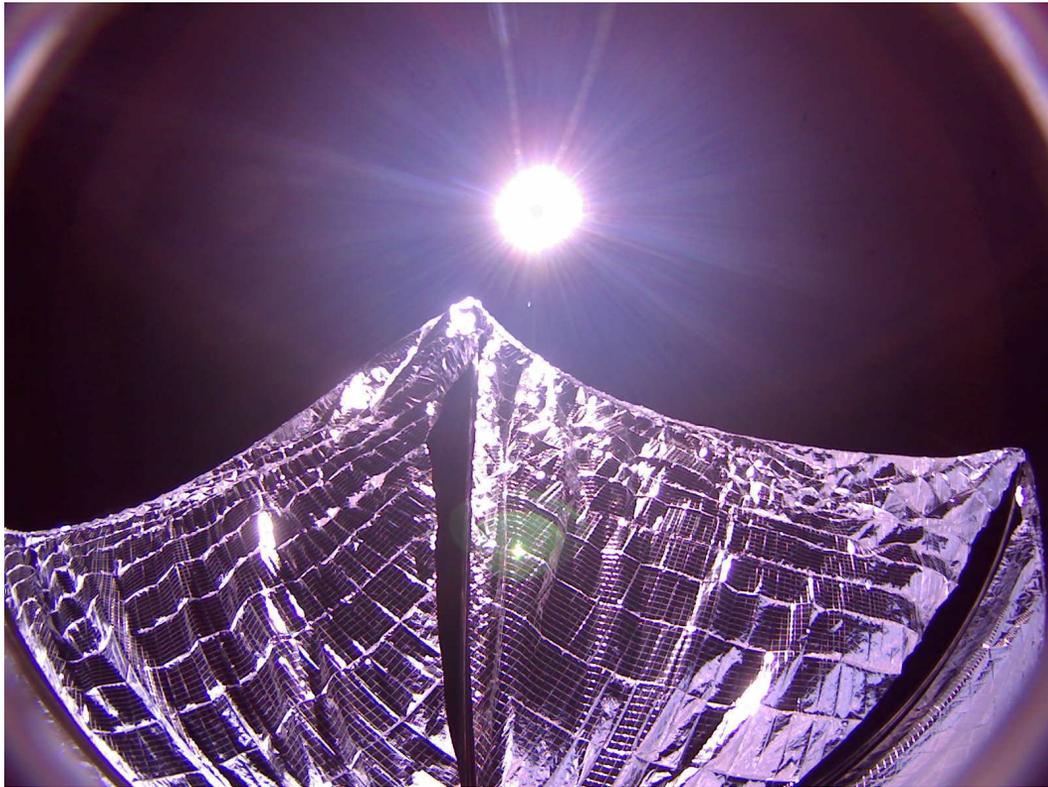
NanoSail-D2 - NASA - Second try launched Nov 19, 2010 Mission complete - 240 days (tumbling) in Near Earth Orbit

Area = 10 m²



LightSail 1 and 2 – Launched by the Planetary Society in 2015 and 2019

Area = 32 m²

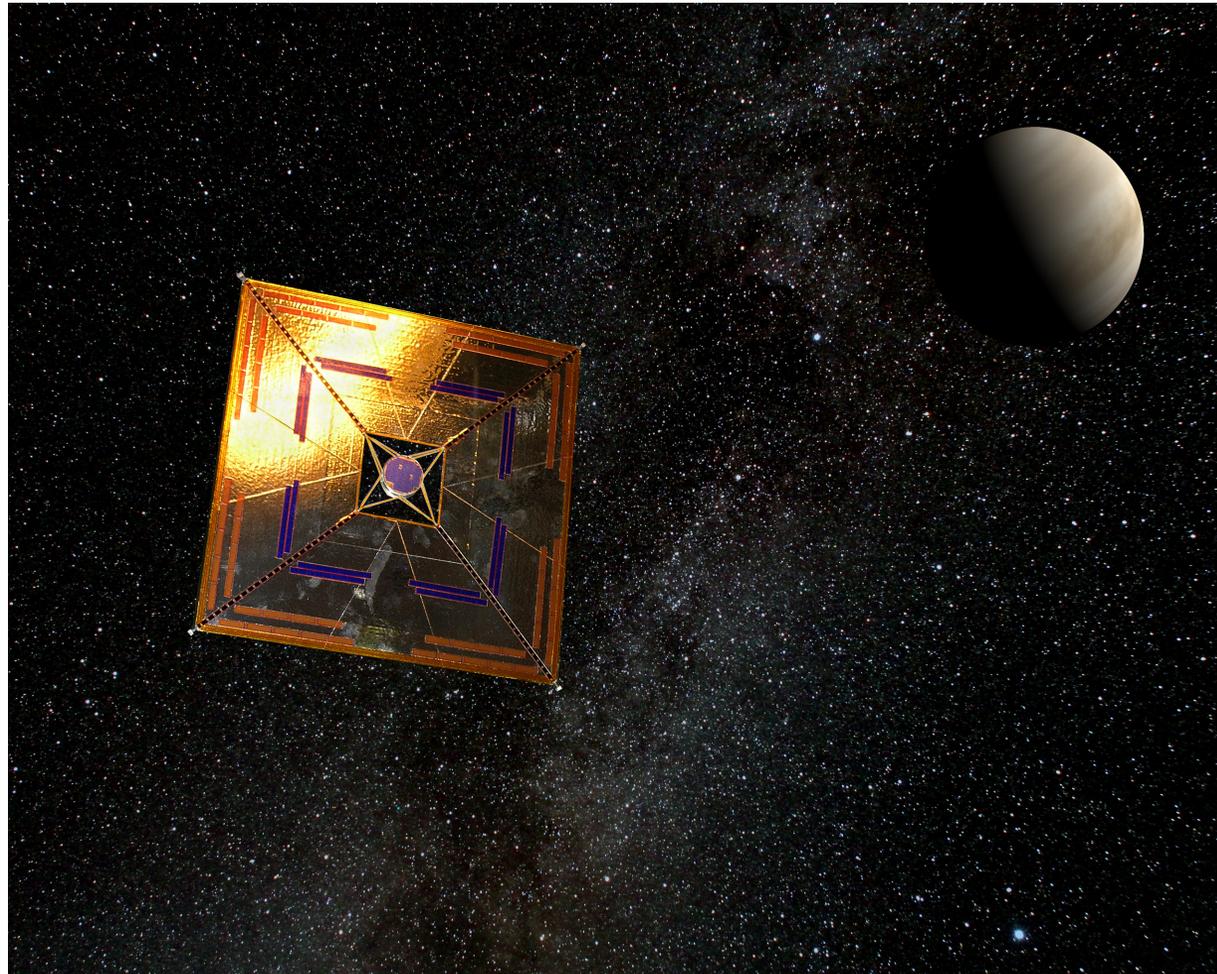


The sails are made of Mylar

IKAROS – Launched by JAXA in May 2010 reached Venus in Dec 2010
First successful planetary mission using a solar sail!

Artist rendering of IKAROS

Sail Area 196 m²,
Thrust 1.12 milliNewtons



IKAROS – Launched by JAXA in 2010 reached Venus in 2015 First successful planetary mission using a solar sail!

IKAROS Mission

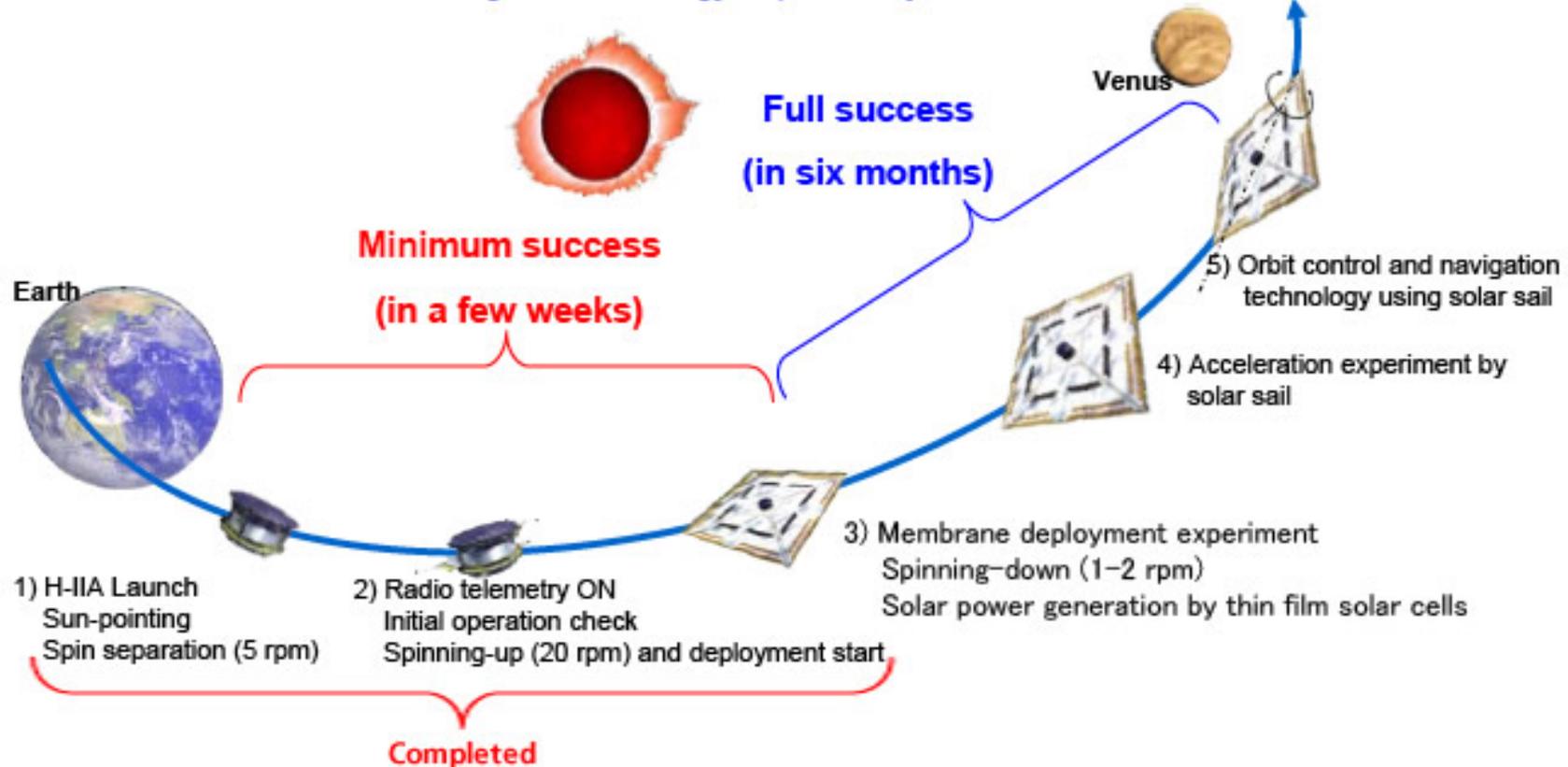
Sail Area 196 m²,
Thrust 1.12 milliNewtons

Minimum success:

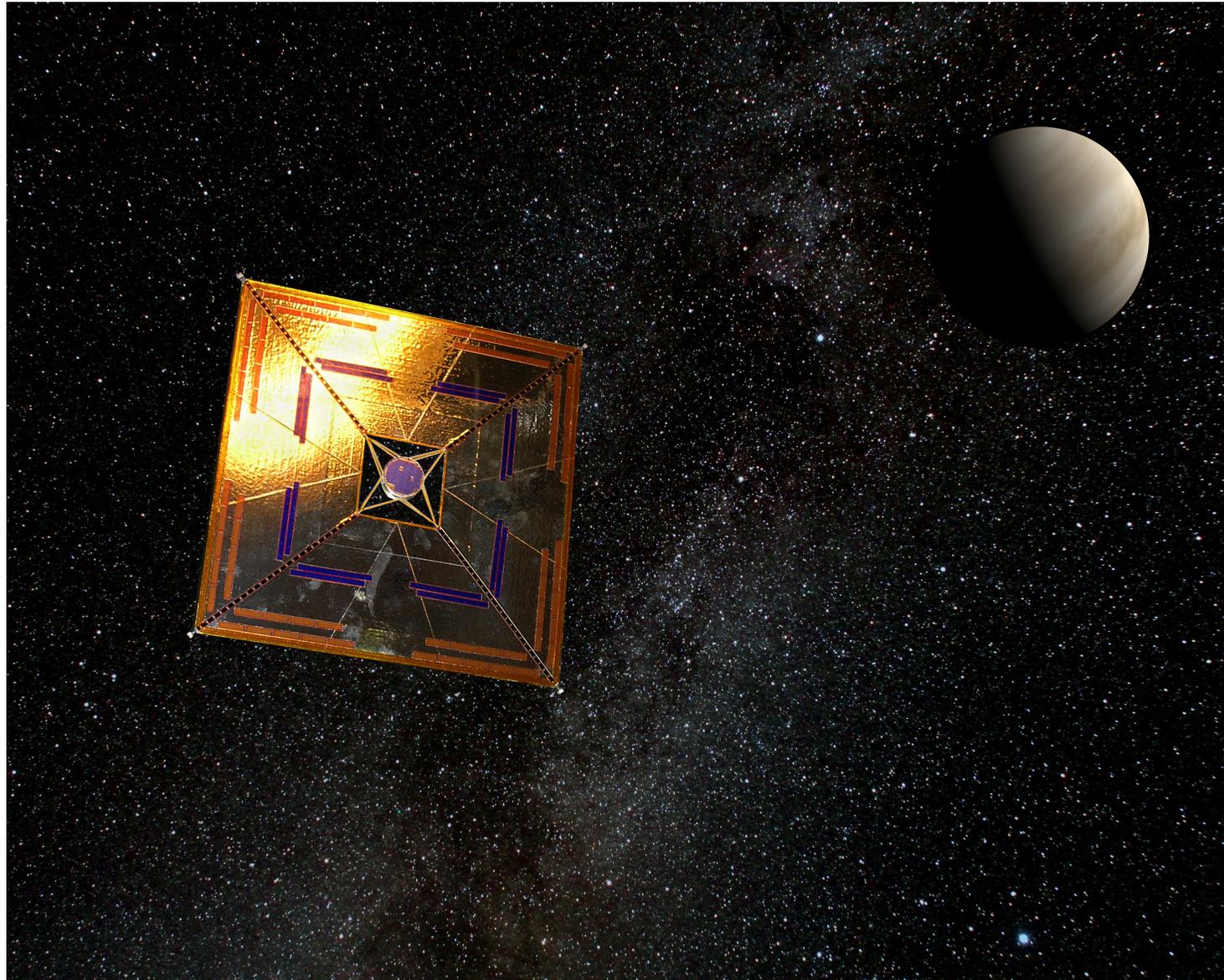
Deployment of the large membrane and power generation by the thin film solar cells.

Full success:

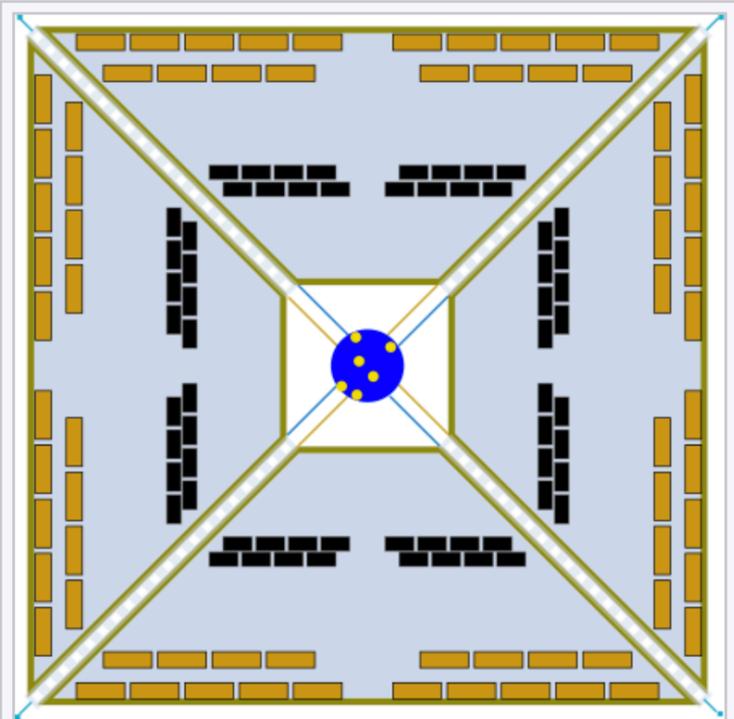
Acceleration verification and navigation technology acquisition by the solar sail.



Artist rendering of IKAROS



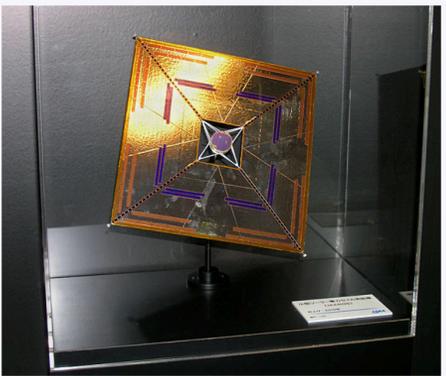
Differential absorption of light by the liquid crystal surfaces used for control.



Sun-facing IKAROS diagram without key

- 1 Tip mass 0.5 kg
- 2 Liquid crystal device
- 3 Membrane 7.5 um thick
- 4 Solar cells 25 um thick
- 5 Tethers
- 6 Main body
- 7 Instruments

IKAROS

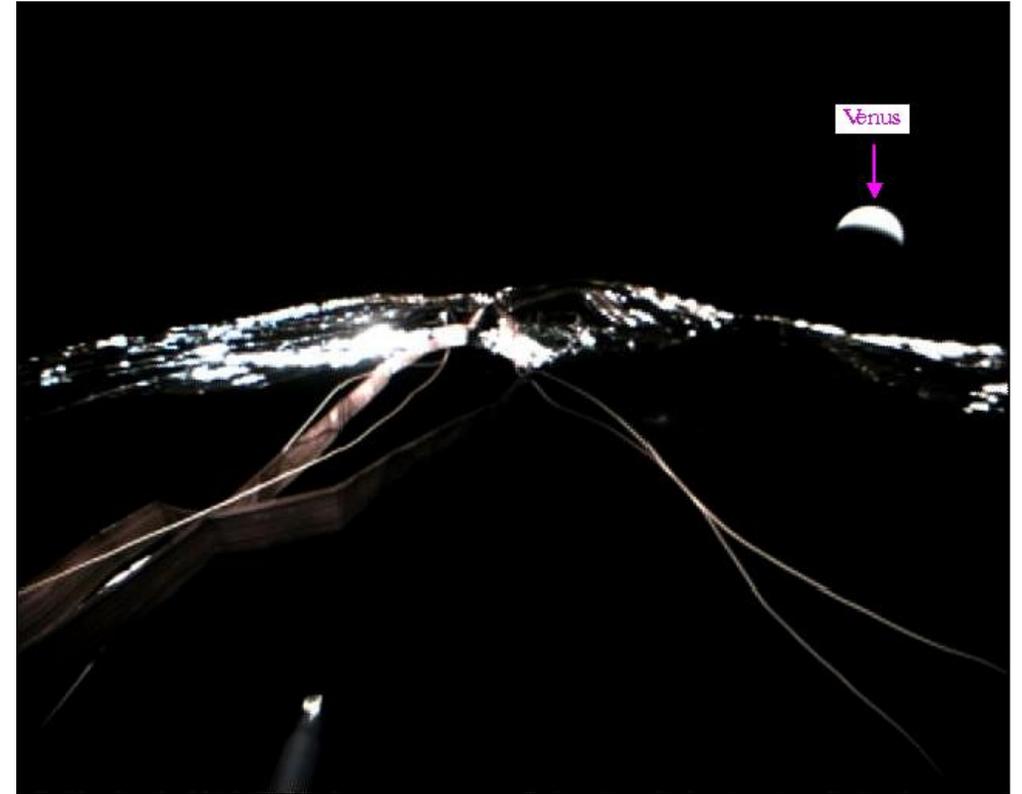
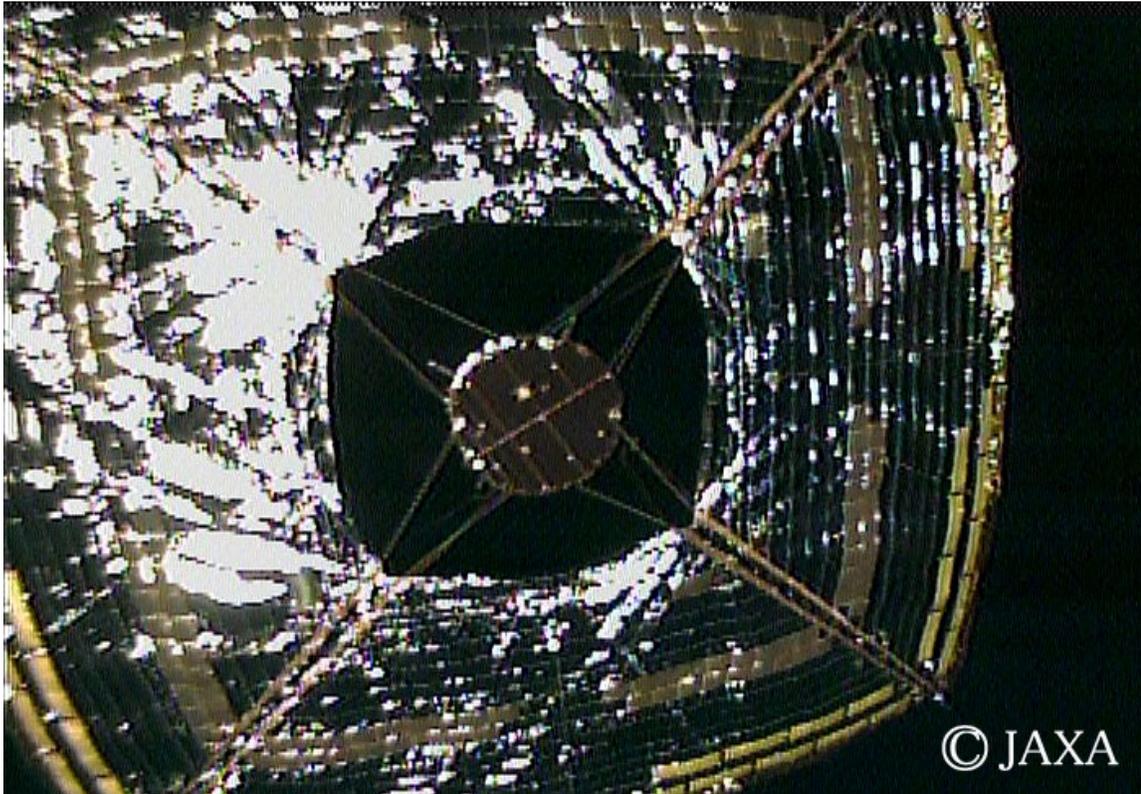


A 1:64 scale model of the IKAROS spacecraft

Mission type	Solar sail technology
Operator	JAXA ^{[1][2][3][4]}
COSPAR ID	2010-020E ↗
SATCAT no.	36577
Website	global.jaxa.jp/projects/sas/ikaros/ ↗
Mission duration	5 years launch to last contact in 2015

Spacecraft properties	
Launch mass	315 kg (694 lb)
Dimensions	Solar sail: 14 m × 14 m (46 ft × 46 ft) (area: 196 m ² (2,110 sq ft)) ^[5]
Start of mission	
Launch date	21:58:22, 20 May 2010 (UTC)
Rocket	H-IIA 202
Launch site	Tanegashima, LA-Y
End of mission	
Last contact	20 May 2015 ^[6]
Orbital parameters	
Reference system	Heliocentric orbit
Flyby of Venus	
Closest approach	8 December 2010
Distance	80,800 kilometers (50,200 mi)

Self images of IKAROS in space

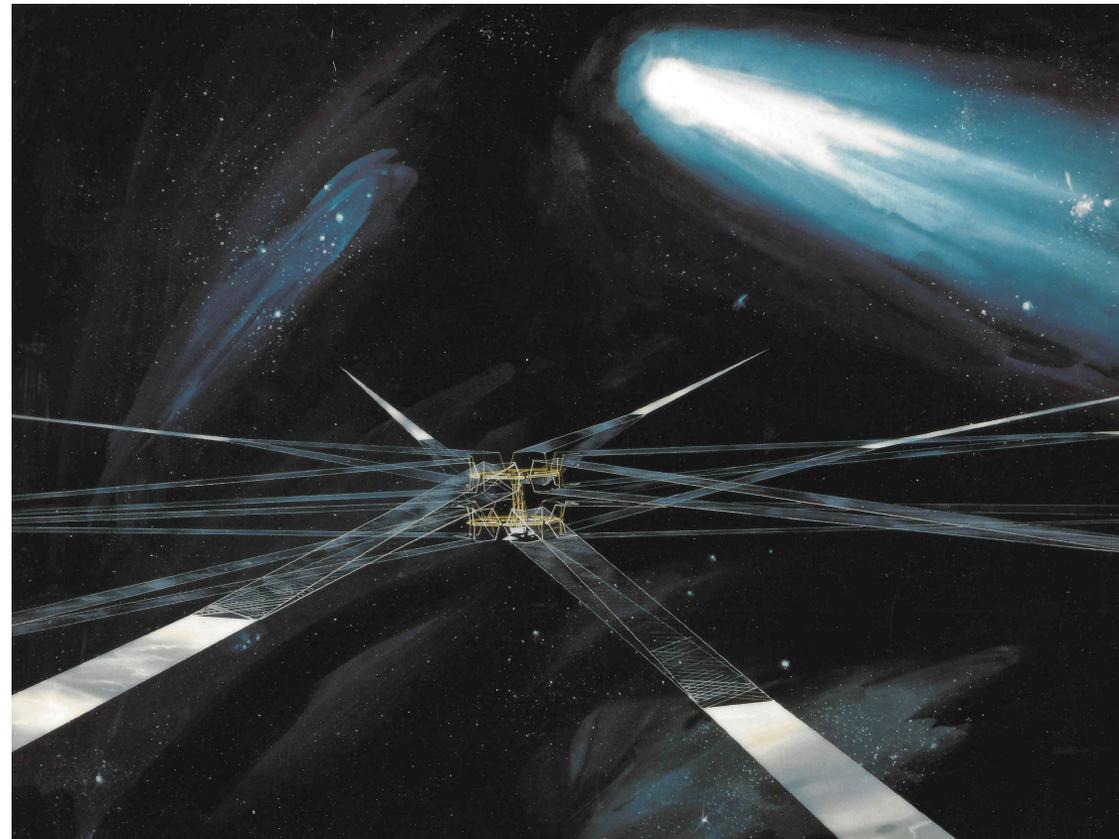


Several Mission Concepts

CubeSail, 20 m² launched on an Electron launch vehicle December 16, 2018 as a first step toward UltraSail

Artist's concept of a heliogyro, proposed to visit Halley's Comet in 1986. Each blade would be 8 m (26 ft) wide and 6.2 km (3.9 mi), for 0.6 km² (0.23 sq mi) of sail area.

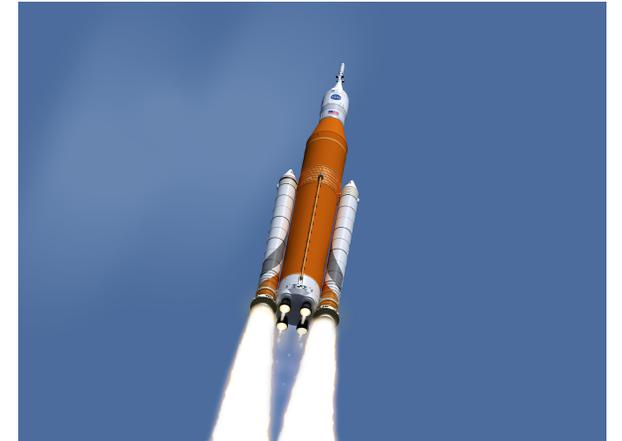
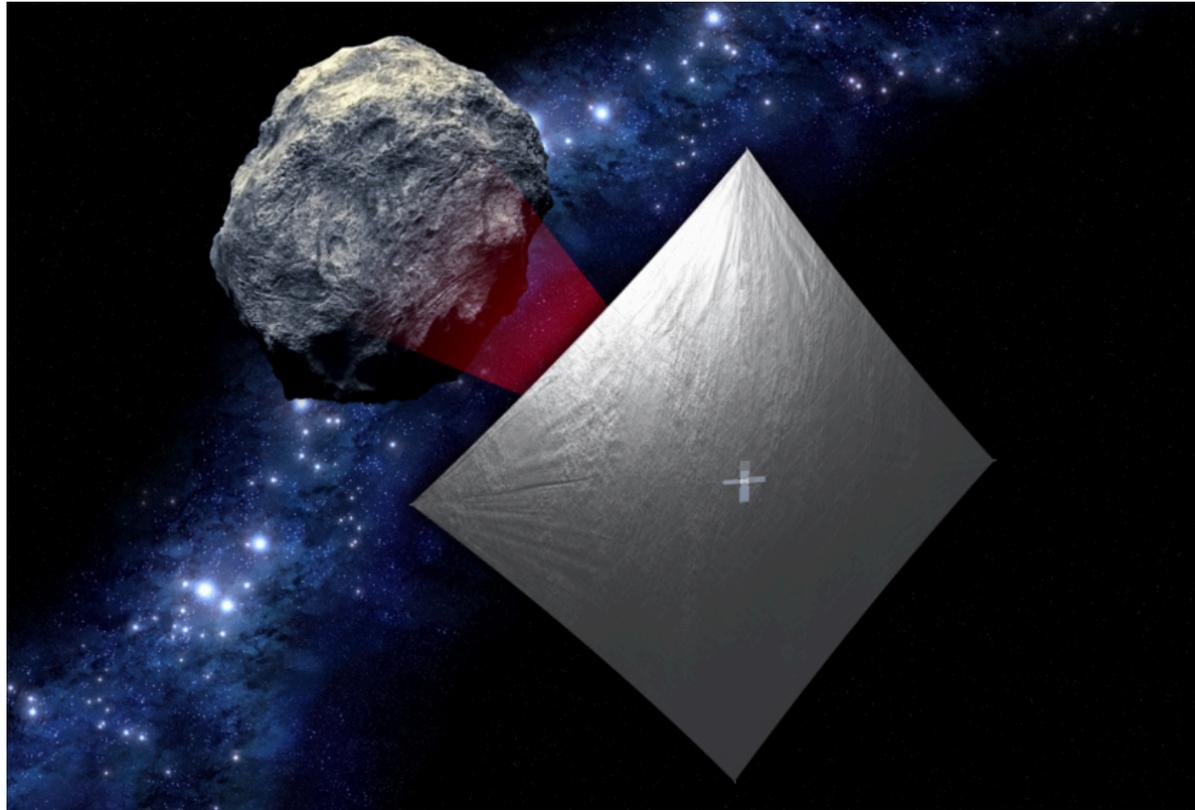
Each blade made of a polyimide film, eg. Kapton, coated with ripstop.



Was CubeSail Successful??

NEA Scout – SLS Pathfinder Launch date Nov 2021

Was supposed to launch on SLS in 2018.



Typical mission times - Earth to Moon

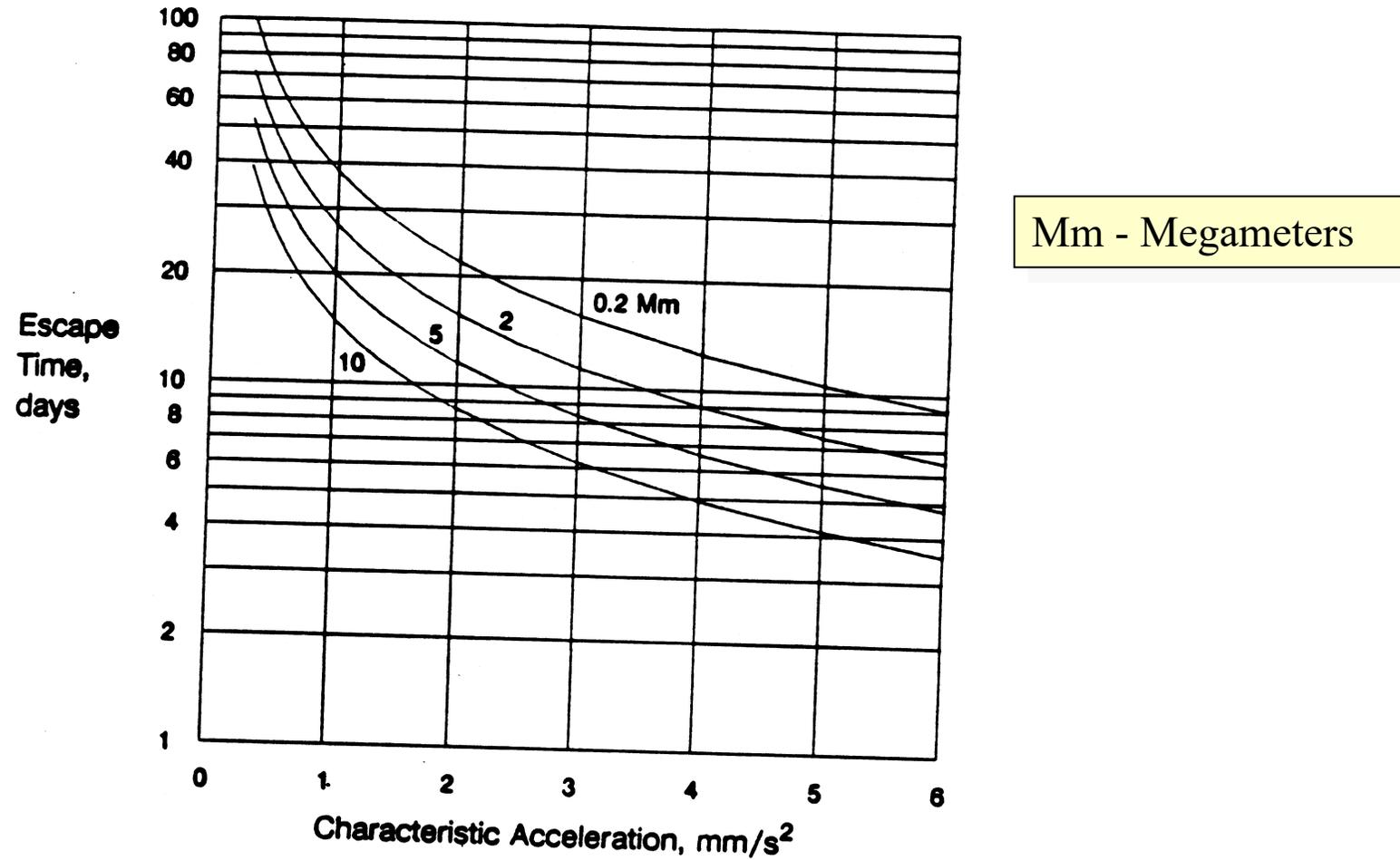
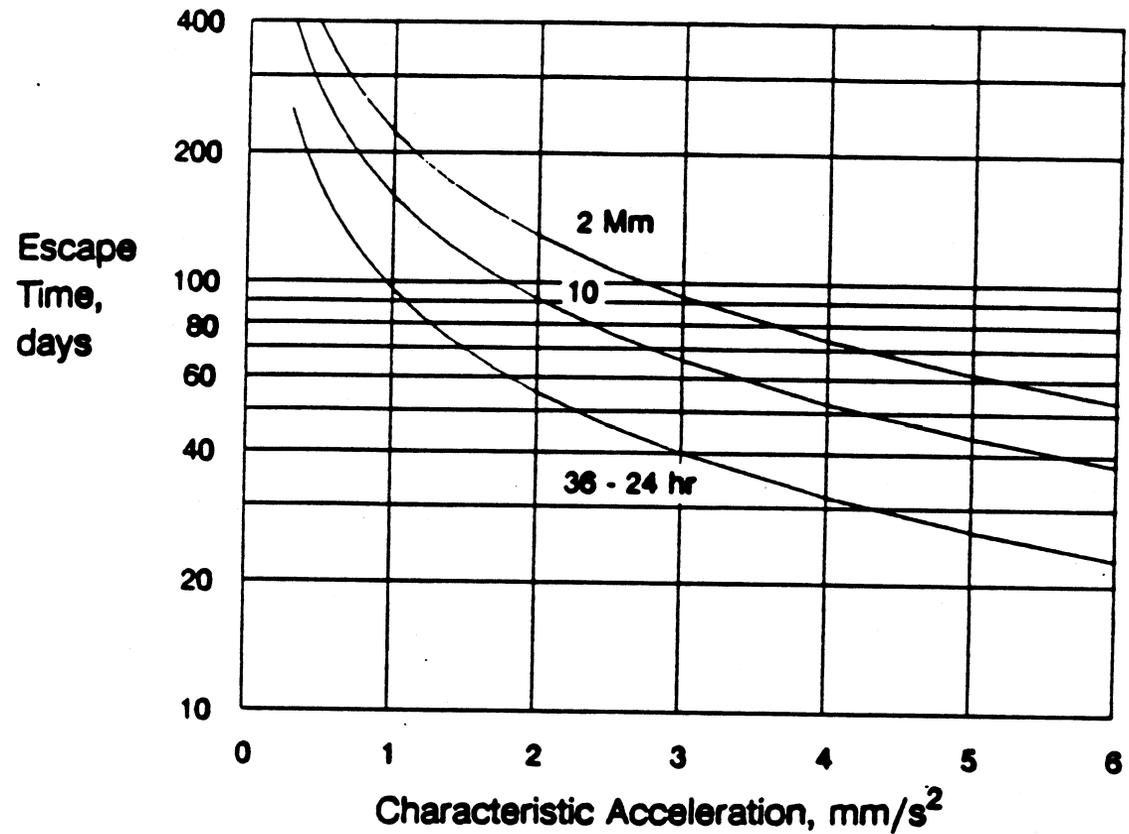


FIGURE 2.2 Lunar Spiral Times.

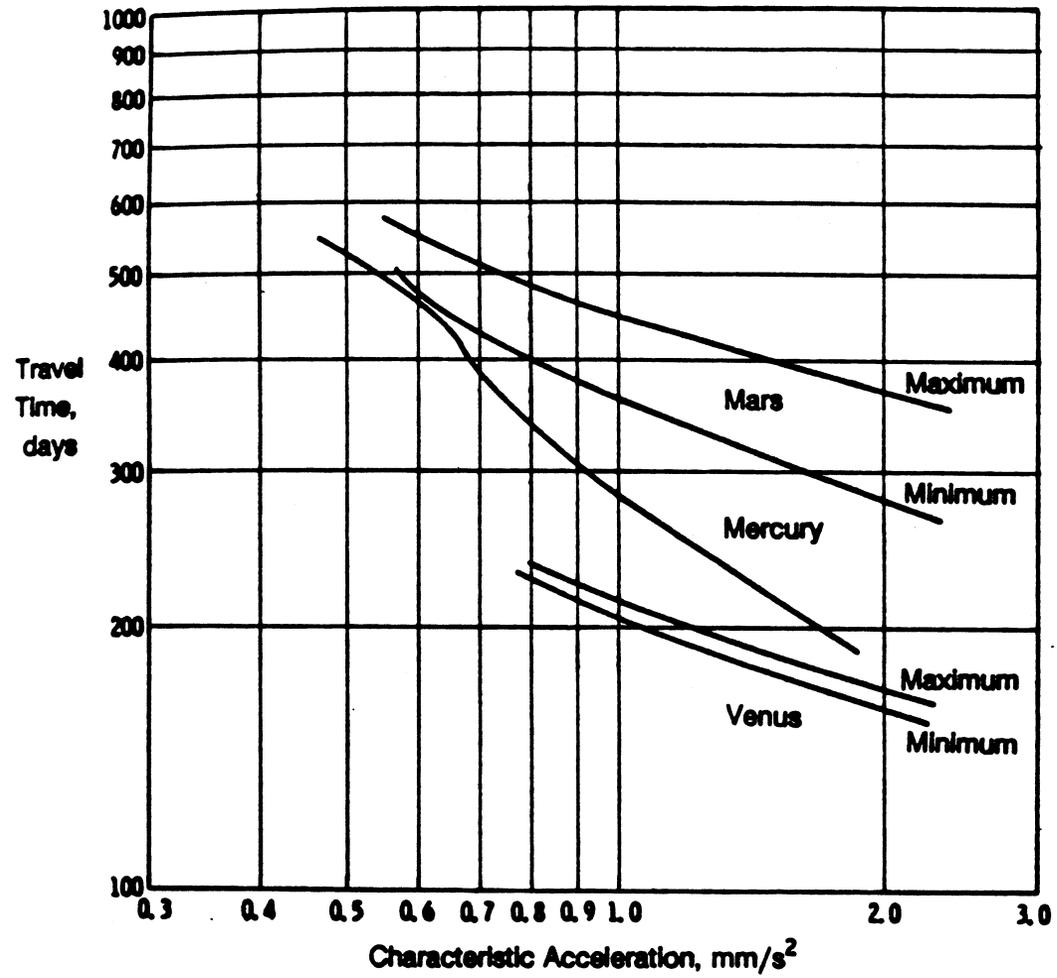
Typical mission times - Earth Escape



Mm - Megameters

FIGURE 2.1 Earth Escape Times From Various Orbits.

Typical mission times - Missions to the Planets

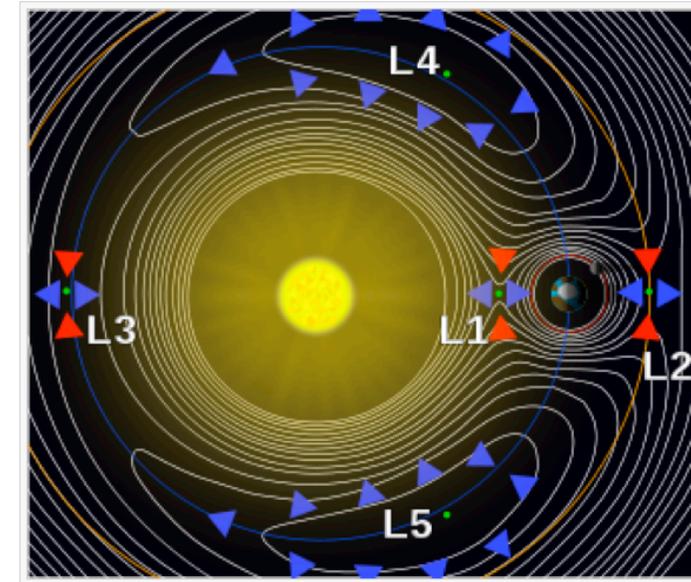
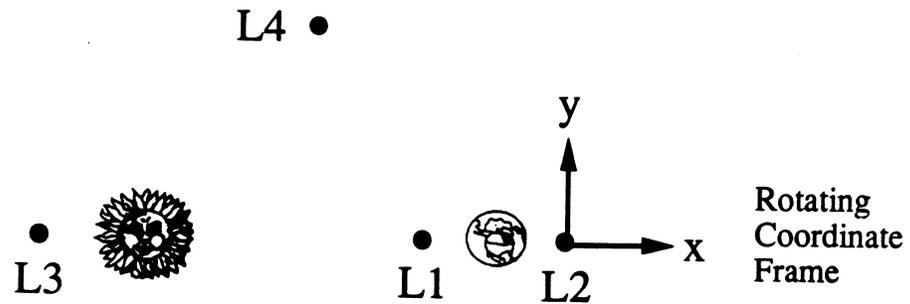


Mm - Megameters

FIGURE 2.4 Typical Travel Times to the Inner Planets.

Mission to L2 – PhD thesis Sun Hur 1992

What is L2 Libration Point?

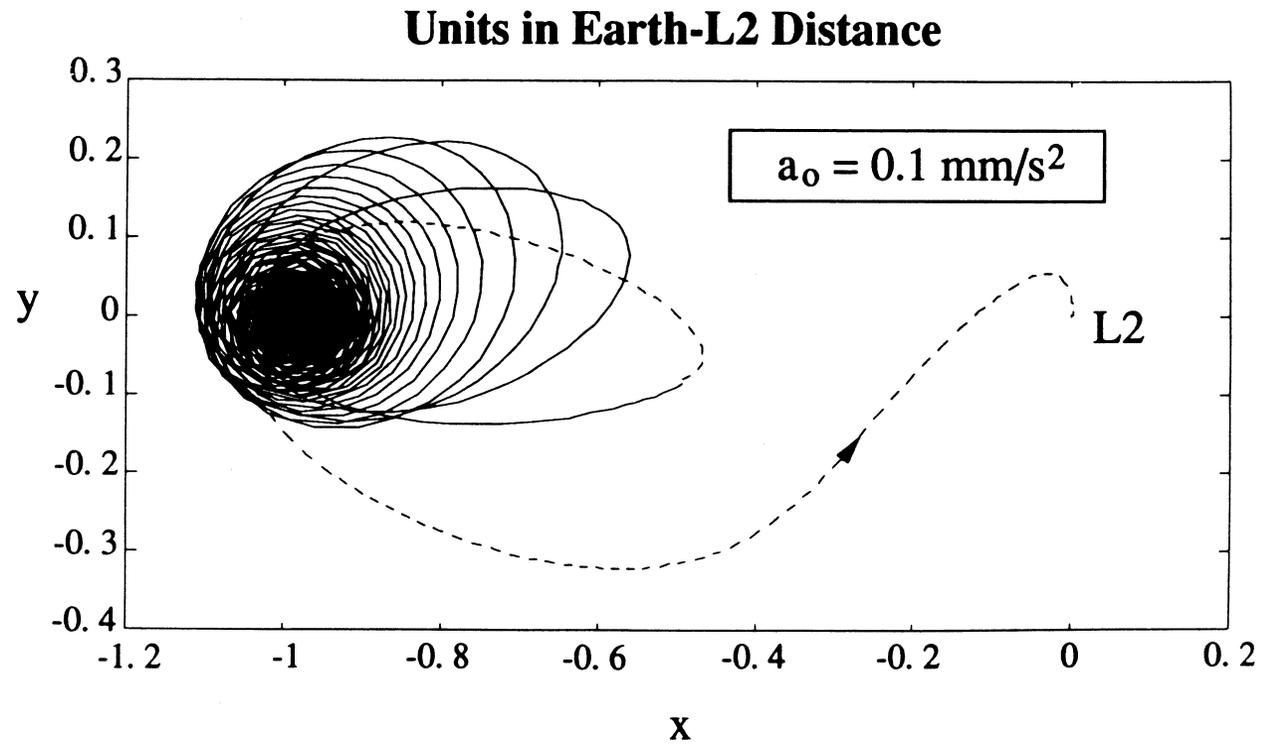


Gravitational potential

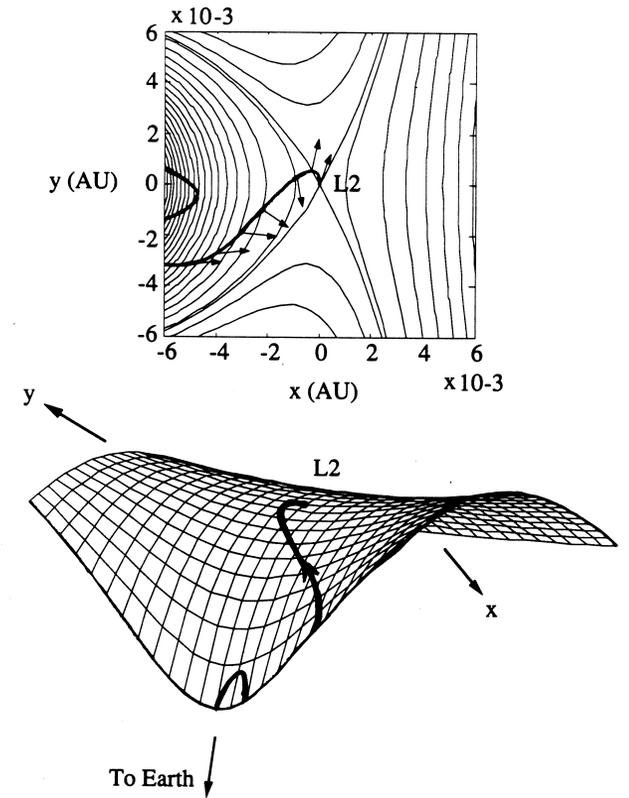
The L2 point is one of five equilibrium points the rotating Sun-Earth system where the gravitational force equals the centrifugal force is an unstable equilibrium point.

CENTRIFUGAL = GRAVITATIONAL

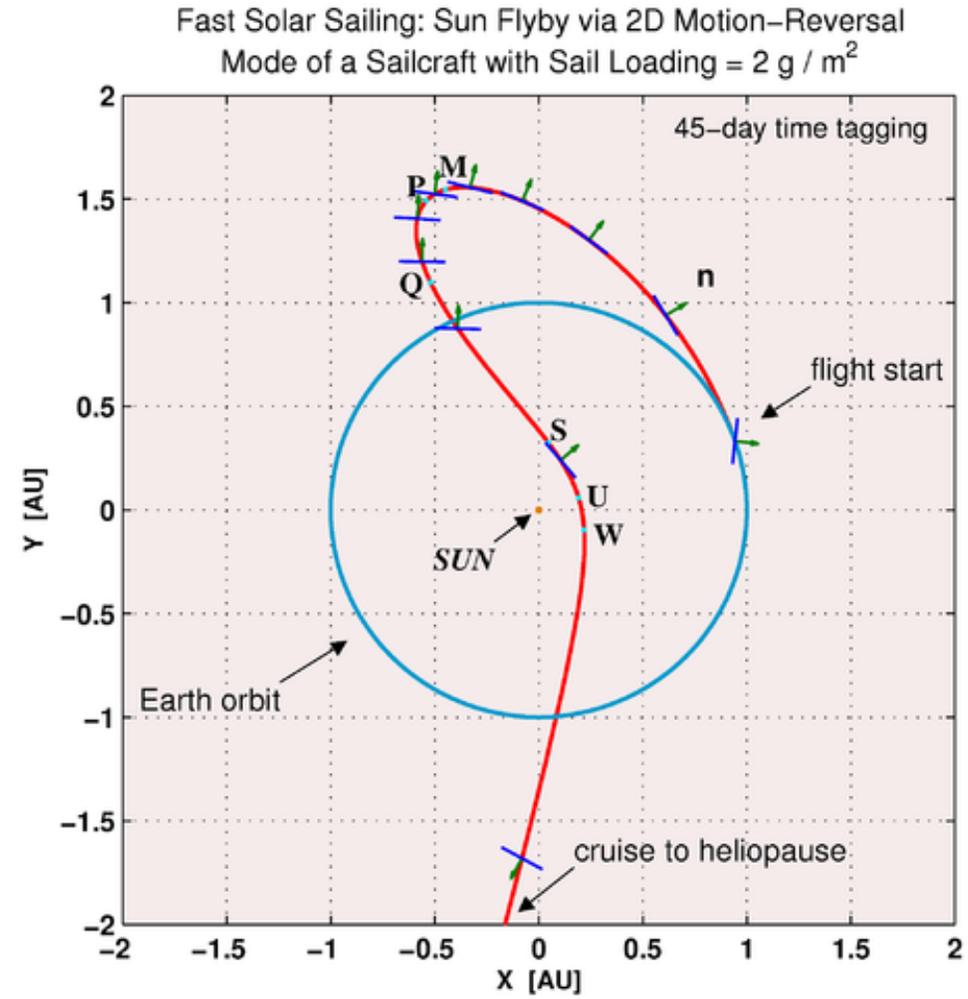
Feasible Trajectory



Trajectory in the modified potential well



Mission Toward the Sun



Mission example - solar watchers

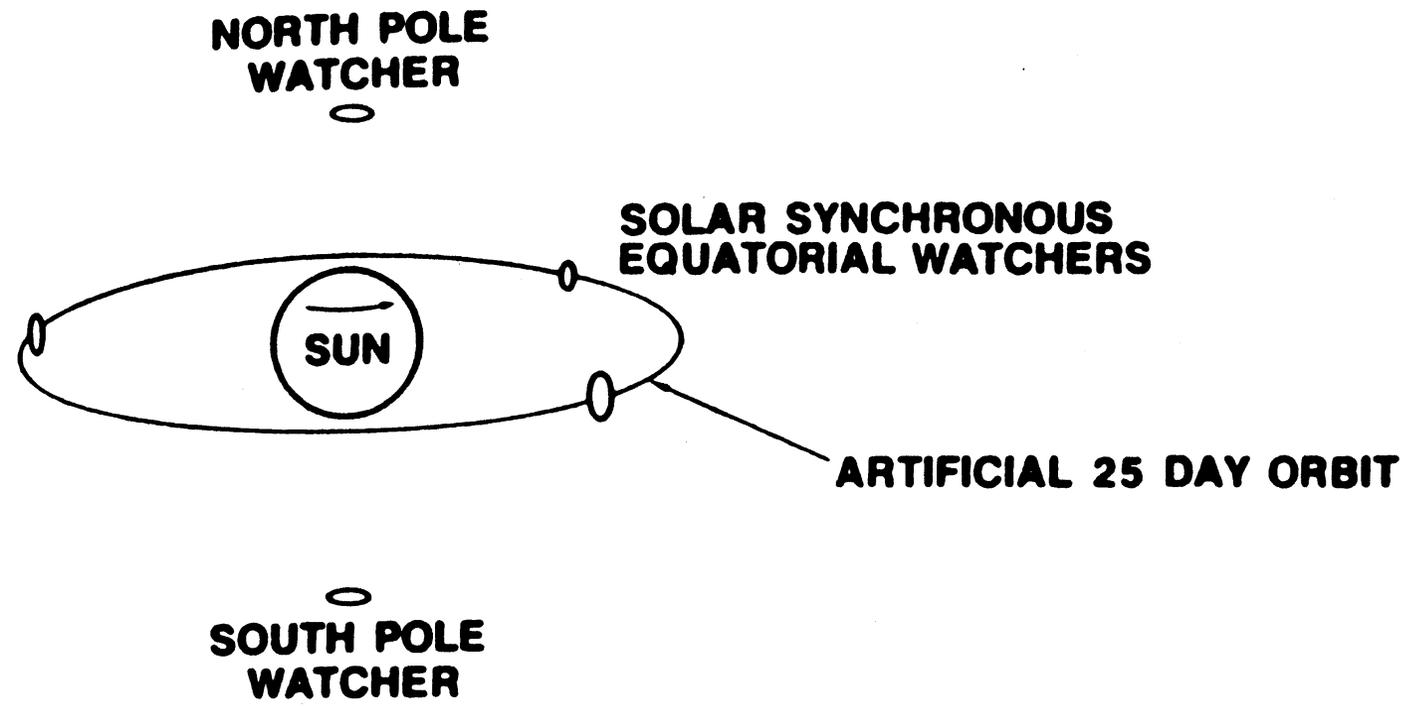


FIGURE 2.21 Synchronous Solar Orbits. (R.L. Forward/Hughes)

Microwave Thrust

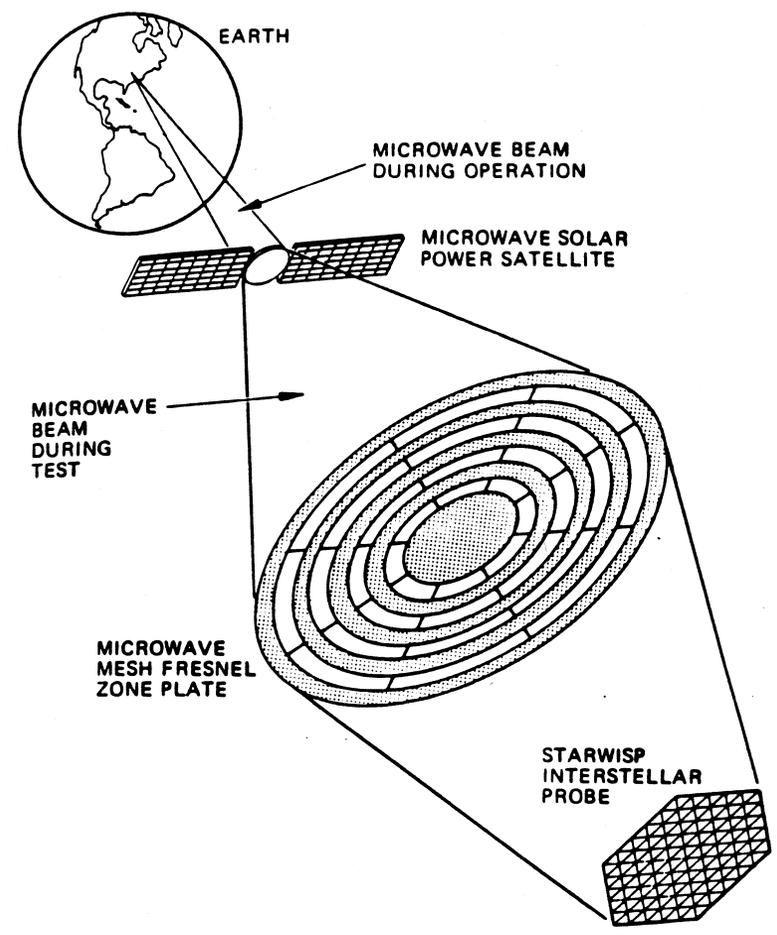


FIGURE 7.10 Operation of a Starwisp Probe. (R.L. Forward)

Mission example - interstellar fly-by

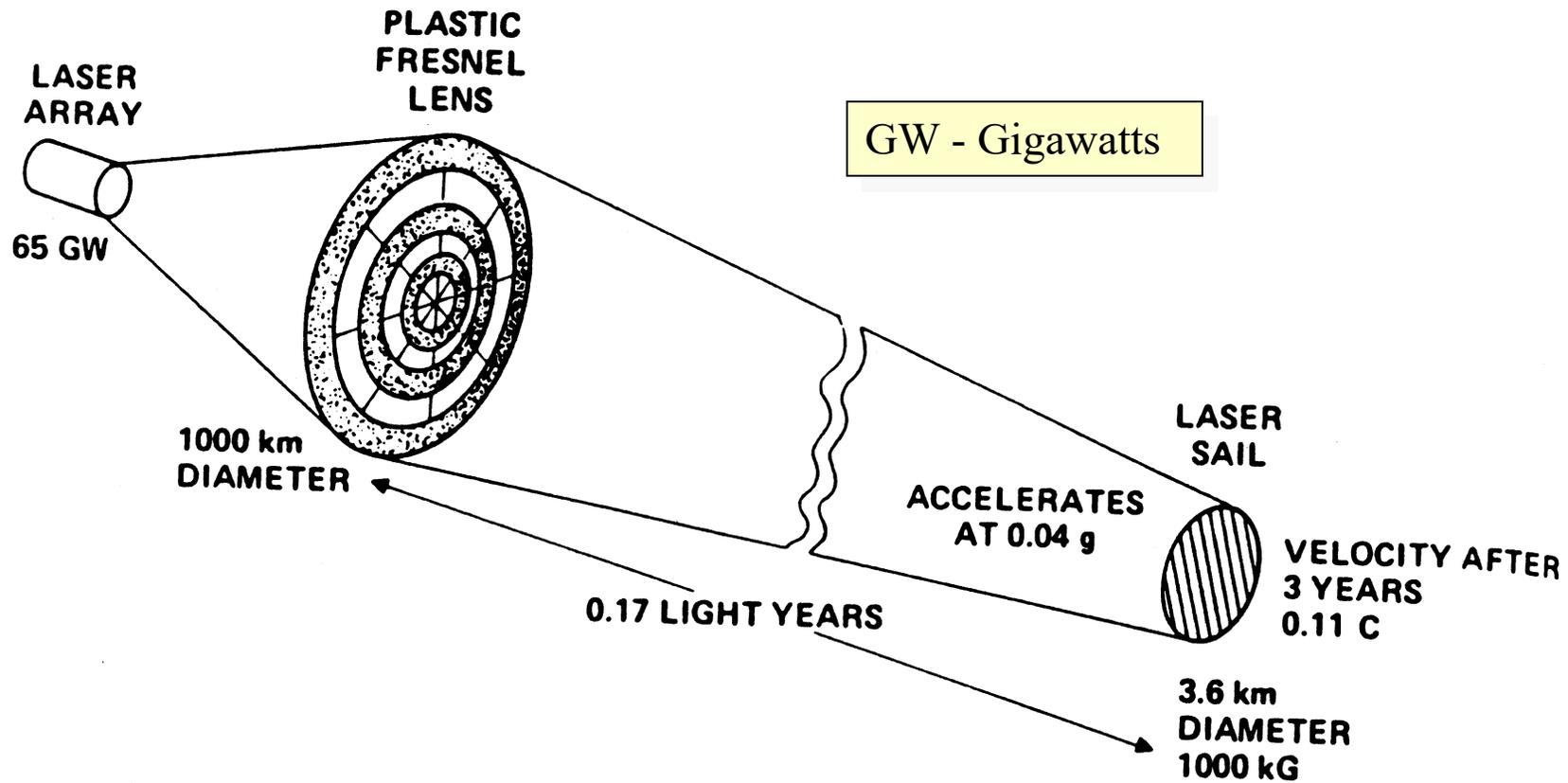
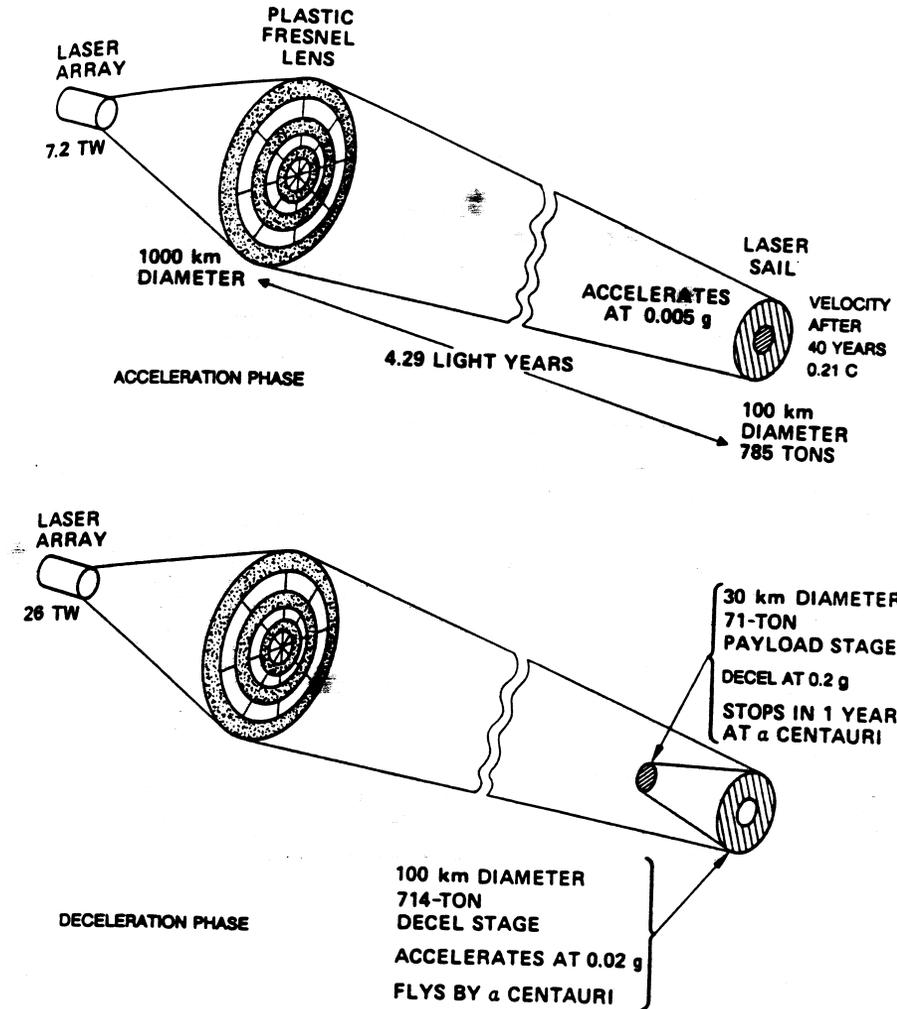


FIGURE 7.6 Profile of an Interstellar Fly-By Probe. (R.L. Forward)

Mission example - one-way interstellar flight



TW - Terawatts

FIGURE 7.7 Profile of a Voyage to Alpha Centauri. (R.L. Forward)

Mission example round-trip interstellar flight

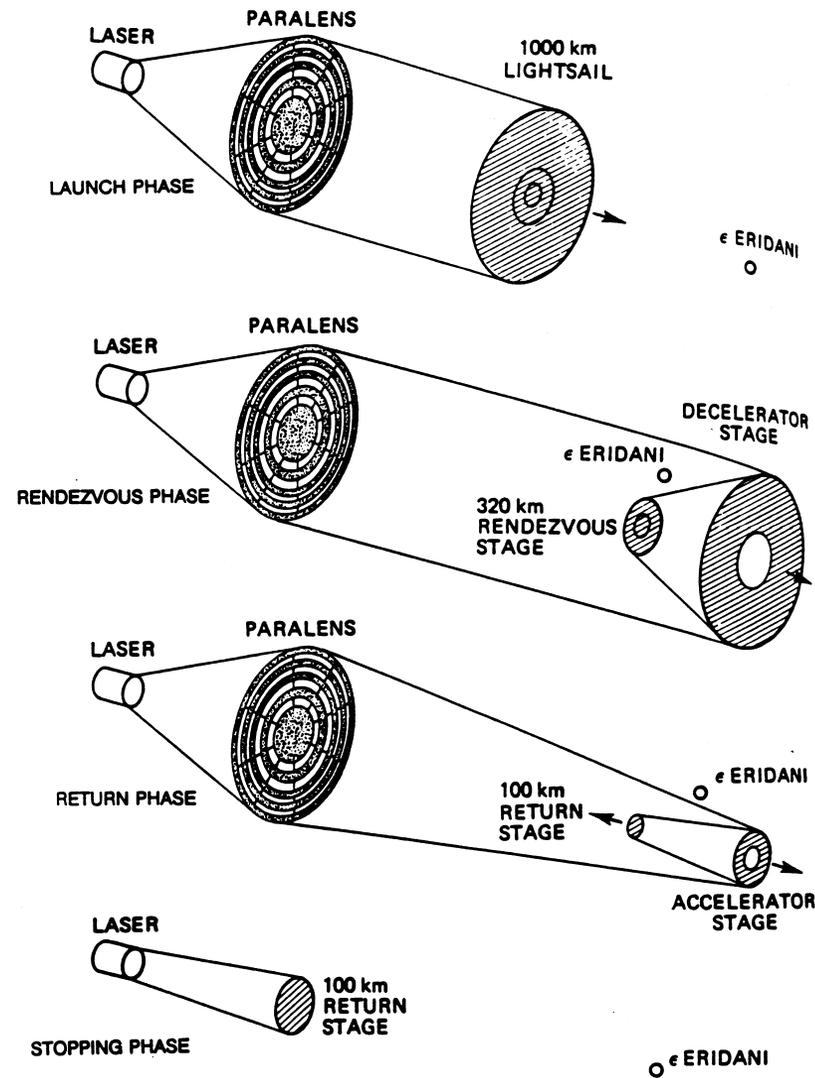
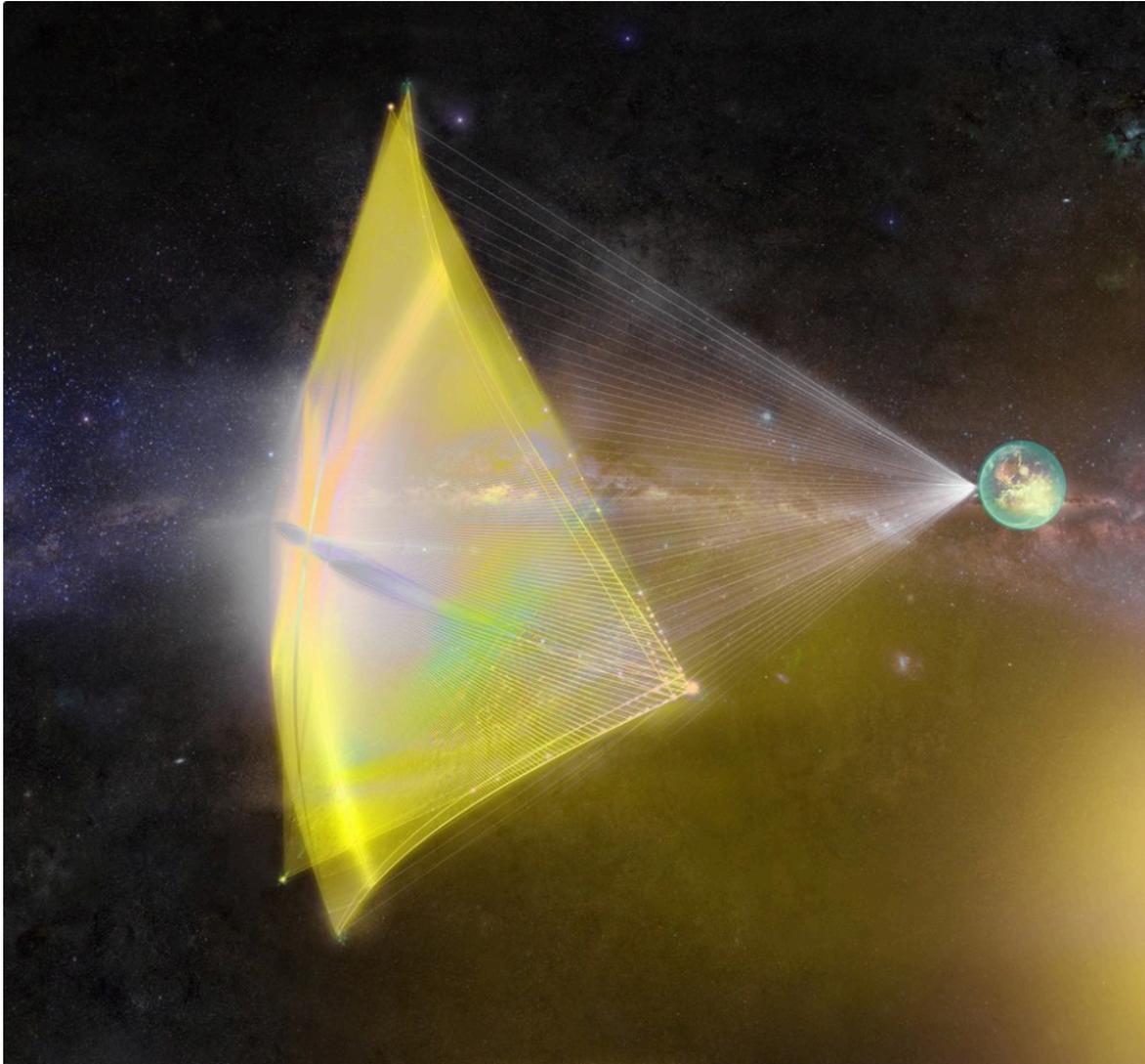


FIGURE 7.8 Profile of a Roundtrip Voyage to Epsilon Eridani. (R.L. Forward)

Breakthrough – Starshot - 2018



Use a Gigawatt scale laser to accelerate a 1 gram nano-sized spacecraft to 20% of the speed of light.

Reach the Alpha Centauri system 4.37 light years away in 20 years.

The *StarChip* – camera, photon thruster, power, navigation and communication.

The *LightSail* – several meters in diameter, only a few hundred atoms thick.

The *LightBeamer* – 100 Gigawatt laser tuned to maximize reflectivity from the sail.

The plan

Path to the stars

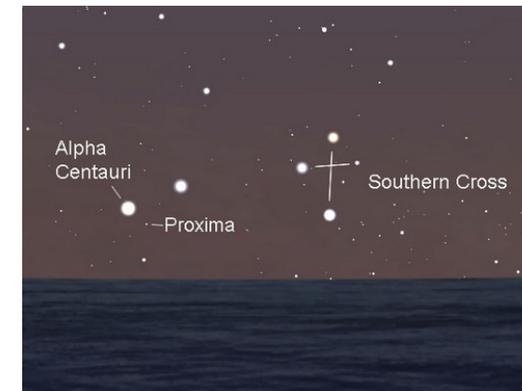
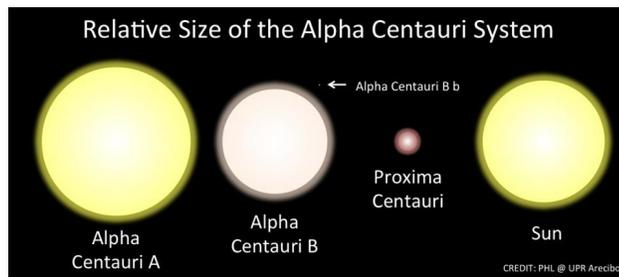
The research and engineering phase is expected to last a number of years. Following that, development of the ultimate mission to Alpha Centauri would require a budget comparable to the **largest current scientific experiments**, and would involve:

- Building a ground-based **kilometer-scale** light beamer at high altitude in dry conditions
- Generating and storing a **few gigawatt hours** of energy per launch
- Launching a 'mothership' carrying **thousands** of nanocrafts to a high-altitude orbit
- Taking advantage of **adaptive optics** technology in real time to compensate for atmospheric effects
- **Focusing** the light beam on the lightsail to accelerate individual nanocrafts to the target speed within minutes
- Accounting for **interstellar dust collisions** en route to the target
- **Capturing images** of a planet, and other scientific data, and transmitting them back to Earth using a compact on-board laser communications system
- Using the same light beamer that launched the nanocrafts to **receive** data from them over 4 years later.

Potential Planets in the Alpha Centauri system

Astronomers estimate that there is a reasonable chance of an **Earth-like planet** existing in the 'habitable zones' of Alpha Centauri's three-star system. A number of scientific instruments, ground-based and space-based, are being developed and enhanced, which will soon identify and characterize planets around nearby stars.

A separate Breakthrough Initiative will support some of these projects.



What would it take to reach the speed of light in a few minutes?

Spacecraft mass = 0.001 kg

Acceleration time to $c/3 = 1000$ sec

Final speed = 10^8 m/sec

Acceleration = 10^5 m/sec²

Force = 10^2 kg-m/sec²

Sail Area = 10 m²

Required light pressure

$$P = \frac{2W}{c} = 10 \text{ N} / \text{m}^2$$

$$W = 1.5 \times 10^9 \text{ J} / \text{m}^2 - \text{sec}$$

The *LightBeamer* – 100 Gigawatt laser tuned to maximize reflectivity from the sail.

How much power does the sail have to dissipate? Assume 99.9999% of the incident energy is reflected by the sail. Only one part in a million is absorbed by the sail.

$$W_{absorbed} = 1.5 \times 10^3 J / m^2 - sec$$

$$Power_{absorbed} = 10 \times W_{absorbed} = 1.5 \times 10^4 J / sec$$

How fast will the sail heat up? Assume the heat capacity of water (very conservative – for metals C is much lower). C=4.184 J/gram-K

$$Power_{absorbed} = 1.5 \times 10^4 J / sec = Cm \frac{dT}{dt}$$

$$C = 4.184 \times 10^3 J / kg - K$$

$$m = 10^{-3} kg$$

$$\frac{dT}{dt} = 3.58 \times 10^2 K / sec$$

How far away is the spacecraft at the end of the acceleration?

$$r = a \frac{t^2}{2} = 0.5 \times 10^5 \times 10^6 = 5 \times 10^{10} m = 0.33 AU$$

Drag caused by the cosmic microwave background

$$\text{Photons per unit volume} = 10^9 \text{ photons/m}^3$$

$$\text{Peak wavelength} = 10^{-6} \text{ m}$$

$$\text{Light sail final speed} = 10^8 \text{ m/sec}$$

$$\text{Photon flux} = 10^{17} \text{ photons/m}^2 \text{ - sec}$$

$$\text{Energy per photon} = hc/\lambda = 6.6 \times 10^{-34} \times 3 \times 10^8 / 10^{-6} = 2 \times 10^{-16} \text{ J}$$

$$\text{Momentum per photon} = h/\lambda = 6.6 \times 10^{-34} / 10^{-6} = 6.6 \times 10^{-28} \text{ kg - m / sec}$$

$$\text{Drag force on a ten square meter sail} = 6.6 \times 10^{-28} \times 10^{17} \times 10 = 6.6 \times 10^{-10} \text{ N}$$

$$\text{Acceleration of a one gram light sail} = 6.6 \times 10^{-7} \text{ m / sec}^2$$

$$\Delta V \text{ over 1000 sec} = 6.6 \times 10^{-4} \text{ m / sec}$$

$$\Delta V \text{ over 20 years} = 20 \times 365 \times 24 \times 3600 \times 6.6 \times 10^{-7} \text{ m / sec} = 416 \text{ m / sec}$$

If the spacecraft speed approached the speed of light this would become an important drag mechanism

This does not account for the blue-shift in the wavelength of the background in the frame of reference of the spacecraft

